

Azimuthal Anisotropy of π^0 Production in Au + Au Collisions at $\sqrt{s_{NN}} = 200$ GeV: Path-Length Dependence of Jet Quenching and the Role of Initial Geometry

A. Adare,¹¹ S. Afanasiev,²⁶ C. Aidala,³⁹ N. N. Ajitanand,⁵⁶ Y. Akiba,^{50,51} H. Al-Bataineh,⁴⁵ J. Alexander,⁵⁶ K. Aoki,^{32,50} Y. Aramaki,¹⁰ E. T. Atomssa,³³ R. Averbeck,⁵⁷ T. C. Awes,⁴⁶ B. Azmoun,⁵ V. Babintsev,²² M. Bai,⁴ G. Baksay,¹⁸ L. Baksay,¹⁸ K. N. Barish,⁶ B. Bassalleck,⁴⁴ A. T. Basye,¹ S. Bathe,⁶ V. Baublis,⁴⁹ C. Baumann,⁴⁰ A. Bazilevsky,⁵ S. Belikov,^{5,*} R. Belmont,⁶¹ R. Bennett,⁵⁷ A. Berdnikov,⁵³ Y. Berdnikov,⁵³ A. A. Bickley,¹¹ J. S. Bok,⁶⁴ K. Boyle,⁵⁷ M. L. Brooks,³⁵ H. Buesching,⁵ V. Bumazhnov,²² G. Bunce,^{5,51} S. Butsyk,³⁵ C. M. Camacho,³⁵ S. Campbell,⁵⁷ C.-H. Chen,⁵⁷ C. Y. Chi,¹² M. Chiu,⁵ I. J. Choi,⁶⁴ R. K. Choudhury,³ P. Christiansen,³⁷ T. Chujo,⁶⁰ P. Chung,⁵⁶ O. Chvala,⁶ V. Cianciolo,⁴⁶ Z. Citron,⁵⁷ B. A. Cole,¹² M. Connors,⁵⁷ P. Constantin,³⁵ M. Csanád,¹⁶ T. Csörgő,²⁹ T. Dahms,⁵⁷ S. Dairaku,^{32,50} I. Danchev,⁶¹ K. Das,¹⁹ A. Datta,³⁹ G. David,⁵ A. Denisov,²² A. Deshpande,^{51,57} E. J. Desmond,⁵ O. Dietzsch,⁵⁴ A. Dion,⁵⁷ M. Donadelli,⁵⁴ O. Drapier,³³ A. Drees,⁵⁷ K. A. Drees,⁴ J. M. Durham,⁵⁷ A. Durum,²² D. Dutta,³ S. Edwards,¹⁹ Y. V. Efremenko,⁴⁶ F. Ellinghaus,¹¹ T. Engelmöre,¹² A. Enokizono,³⁴ H. En'yo,^{50,51} S. Esumi,⁶⁰ B. Fadem,⁴¹ D. E. Fields,⁴⁴ M. Finger, Jr.,⁷ M. Finger,⁷ F. Fleuret,³³ S. L. Fokin,³¹ Z. Fraenkel,^{63,*} J. E. Frantz,⁵⁷ A. Franz,⁵ A. D. Frawley,¹⁹ K. Fujiwara,⁵⁰ Y. Fukao,⁵⁰ T. Fusayasu,⁴³ I. Garishvili,⁵⁸ A. Glenn,¹¹ H. Gong,⁵⁷ M. Gonin,³³ Y. Goto,^{50,51} R. Granier de Cassagnac,³³ N. Grau,¹² S. V. Greene,⁶¹ M. Grosse Perdekamp,^{23,51} T. Gunji,¹⁰ H.-Å. Gustafsson,^{37,*} J. S. Haggerty,⁵ K. I. Hahn,¹⁷ H. Hamagaki,¹⁰ J. Hamblen,⁵⁸ J. Hanks,¹² R. Han,⁴⁸ E. P. Hartouni,³⁴ E. Haslum,³⁷ R. Hayano,¹⁰ M. Heffner,³⁴ S. Hegyi,²⁹ T. K. Hemmick,⁵⁷ T. Hester,⁶ X. He,²⁰ J. C. Hill,²⁵ M. Hohlmann,¹⁸ W. Holzmann,¹² K. Homma,²¹ B. Hong,³⁰ T. Horaguchi,²¹ D. Hornback,⁵⁸ S. Huang,⁶¹ T. Ichihara,^{50,51} R. Ichimiya,⁵⁰ J. Ide,⁴¹ Y. Ikeda,⁶⁰ K. Imai,^{32,50} M. Inaba,⁶⁰ D. Isenhower,¹ M. Ishihara,⁵⁰ T. Isobe,¹⁰ M. Issah,⁶¹ A. Isupov,²⁶ D. Ivanischev,⁴⁹ B. V. Jacak,^{57,†} J. Jia,^{5,56} J. Jin,¹² B. M. Johnson,⁵ K. S. Joo,⁴² D. Jouan,⁴⁷ D. S. Jumper,¹ F. Kajihara,¹⁰ S. Kametani,⁵⁰ N. Kamihara,⁵¹ J. Kamin,⁵⁷ J. H. Kang,⁶⁴ J. Kapustinsky,³⁵ K. Karatsu,³² D. Kallow,^{39,51} M. Kawashima,^{52,50} A. V. Kazantsev,³¹ T. Kempel,²⁵ A. Khanzadeev,⁴⁹ K. M. Kijima,²¹ B. I. Kim,³⁰ D. H. Kim,⁴² D. J. Kim,²⁷ E. J. Kim,⁸ E. Kim,⁵⁵ S. H. Kim,⁶⁴ Y. J. Kim,²³ E. Kinney,¹¹ K. Kiriluk,¹¹ Á. Kiss,¹⁶ E. Kistenev,⁵ L. Kochenda,⁴⁹ B. Komkov,⁴⁹ M. Konno,⁶⁰ J. Koster,²³ D. Kotchetkov,⁴⁴ A. Kozlov,⁶³ A. Král,¹³ A. Kravitz,¹² G. J. Kunde,³⁵ K. Kurita,^{52,50} M. Kurosawa,⁵⁰ Y. Kwon,⁶⁴ G. S. Kyle,⁴⁵ R. Lacey,⁵⁶ Y. S. Lai,¹² J. G. Lajoie,²⁵ A. Lebedev,²⁵ D. M. Lee,³⁵ J. Lee,¹⁷ K. B. Lee,³⁰ K. Lee,⁵⁵ K. S. Lee,³⁰ M. J. Leitch,³⁵ M. A. L. Leite,⁵⁴ E. Leitner,⁶¹ B. Lenzi,⁵⁴ P. Liebing,⁵¹ L. A. Linden Levy,¹¹ T. Liška,¹³ A. Litvinenko,²⁶ H. Liu,^{35,45} M. X. Liu,³⁵ X. Li,⁹ B. Love,⁶¹ R. Luechtenborg,⁴⁰ D. Lynch,⁵ C. F. Maguire,⁶¹ Y. I. Makdisi,⁴ A. Malakhov,²⁶ M. D. Malik,⁴⁴ V. I. Manko,³¹ E. Mannel,¹² Y. Mao,^{48,50} H. Masui,⁶⁰ F. Matathias,¹² M. McCumber,⁵⁷ P. L. McGaughey,³⁵ N. Means,⁵⁷ B. Meredith,²³ Y. Miake,⁶⁰ A. C. Mignerey,³⁸ P. Mikeš,^{7,24} K. Miki,⁶⁰ A. Milov,⁵ M. Mishra,² J. T. Mitchell,⁵ A. K. Mohanty,³ Y. Morino,¹⁰ A. Morreale,⁶ D. P. Morrison,⁵ T. V. Moukhanova,³¹ J. Murata,^{52,50} S. Nagamiya,²⁸ J. L. Nagle,¹¹ M. Naglis,⁶³ M. I. Nagy,¹⁶ I. Nakagawa,^{50,51} Y. Nakamiya,²¹ T. Nakamura,^{21,28} K. Nakano,^{50,59} J. Newby,³⁴ M. Nguyen,⁵⁷ R. Nouicer,⁵ A. S. Nyanin,³¹ E. O'Brien,⁵ S. X. Oda,¹⁰ C. A. Ogilvie,²⁵ K. Okada,⁵¹ M. Oka,⁶⁰ Y. Onuki,⁵⁰ A. Oskarsson,³⁷ M. Ouchida,²¹ K. Ozawa,¹⁰ R. Pak,⁵ V. Pantuev,⁵⁷ V. Papavassiliou,⁴⁵ I. H. Park,¹⁷ J. Park,⁵⁵ S. K. Park,³⁰ W. J. Park,³⁰ S. F. Pate,⁴⁵ H. Pei,²⁵ J.-C. Peng,²³ H. Pereira,¹⁴ V. Peresedov,²⁶ D. Yu. Peressounko,³¹ C. Pinkenburg,⁵ R. P. Pisani,⁵ M. Proissl,⁵⁷ M. L. Purschke,⁵ A. K. Purwar,³⁵ H. Qu,²⁰ J. Rak,²⁷ A. Rakotozafindrabe,³³ I. Ravinovich,⁶³ K. F. Read,^{46,58} K. Reygers,⁴⁰ V. Riabov,⁴⁹ Y. Riabov,⁴⁹ E. Richardson,³⁸ D. Roach,⁶¹ G. Roche,³⁶ S. D. Rolnick,⁶ M. Rosati,²⁵ C. A. Rosen,¹¹ S. S. E. Rosendahl,³⁷ P. Rosnet,³⁶ P. Rukoyatkin,²⁶ P. Ružička,²⁴ B. Sahlmueller,⁴⁰ N. Saito,²⁸ T. Sakaguchi,⁵ K. Sakashita,^{50,59} V. Samsonov,⁴⁹ S. Sano,^{10,62} T. Sato,⁶⁰ S. Sawada,²⁸ K. Sedgwick,⁶ J. Seele,¹¹ R. Seidl,²³ A. Yu. Semenov,²⁵ R. Seto,⁶ D. Sharma,⁶³ I. Shein,²² T.-A. Shibata,^{50,59} K. Shigaki,²¹ M. Shimomura,⁶⁰ K. Shoji,^{32,50} P. Shukla,³ A. Sickles,⁵ C. L. Silva,⁵⁴ D. Silvermyr,⁴⁶ C. Silvestre,¹⁴ K. S. Sim,³⁰ B. K. Singh,² C. P. Singh,² V. Singh,² M. Slunečka,⁷ R. A. Soltz,³⁴ W. E. Sondheim,³⁵ S. P. Sorensen,⁵⁸ I. V. Sourikova,⁵ N. A. Sparks,¹ P. W. Stankus,⁴⁶ E. Stenlund,³⁷ S. P. Stoll,⁵ T. Sugitate,²¹ A. Sukhanov,⁵ J. Sziklai,²⁹ E. M. Takagui,⁵⁴ A. Taketani,^{50,51} R. Tanabe,⁶⁰ Y. Tanaka,⁴³ K. Tanida,^{32,50,51,55} M. J. Tannenbaum,⁵ S. Tarafdar,² A. Taranenko,⁵⁶ P. Tarján,¹⁵ H. Themann,⁵⁷ T. L. Thomas,⁴⁴ M. Togawa,^{32,50} A. Toia,⁵⁷ L. Tomášek,²⁴ H. Torii,²¹ R. S. Towell,¹ I. Tserruya,⁶³ Y. Tsuchimoto,²¹ C. Vale,^{5,25} H. Valle,⁶¹ H. W. van Hecke,³⁵ E. Vazquez-Zambrano,¹² A. Veicht,²³ J. Velkovska,⁶¹ R. Vértési,^{15,29} A. A. Vinogradov,³¹ M. Virius,¹³ V. Vrba,²⁴ E. Vznuzdaev,⁴⁹ X. R. Wang,⁴⁵ D. Watanabe,²¹ K. Watanabe,⁶⁰ Y. Watanabe,^{50,51} F. Wei,²⁵ R. Wei,⁵⁶ J. Wessels,⁴⁰ S. N. White,⁵ D. Winter,¹² J. P. Wood,¹ C. L. Woody,⁵ R. M. Wright,¹ M. Wysocki,¹¹ W. Xie,⁵¹ Y. L. Yamaguchi,¹⁰ K. Yamaura,²¹ R. Yang,²³ A. Yanovich,²² J. Ying,²⁰ S. Yokkaichi,^{50,51}

G. R. Young,⁴⁶ I. Younus,⁴⁴ Z. You,⁴⁸ I. E. Yushmanov,³¹ W. A. Zajc,¹² C. Zhang,⁴⁶ S. Zhou,⁹ and L. Zolin²⁶

(PHENIX Collaboration)

- ¹Abilene Christian University, Abilene, Texas 79699, USA
²Department of Physics, Banaras Hindu University, Varanasi 221005, India
³Bhabha Atomic Research Centre, Bombay 400 085, India
⁴Collider-Accelerator Department, Brookhaven National Laboratory, Upton, New York 11973-5000, USA
⁵Physics Department, Brookhaven National Laboratory, Upton, New York 11973-5000, USA
⁶University of California—Riverside, Riverside, California 92521, USA
⁷Charles University, Ovocný trh 5, Praha 1, 116 36, Prague, Czech Republic
⁸Chonbuk National University, Jeonju 561-756, Korea
⁹China Institute of Atomic Energy (CIAE), Beijing, People's Republic of China
¹⁰Center for Nuclear Study, Graduate School of Science, University of Tokyo, 7-3-1 Hongo, Bunkyo, Tokyo 113-0033, Japan
¹¹University of Colorado, Boulder, Colorado 80309, USA
¹²Columbia University, New York, New York 10027 and Nevis Laboratories, Irvington, New York 10533, USA
¹³Czech Technical University, Zikova 4, 166 36 Prague 6, Czech Republic
¹⁴Dapnia, CEA Saclay, F-91191, Gif-sur-Yvette, France
¹⁵Debrecen University, H-4010 Debrecen, Egyetem tér 1, Hungary
¹⁶ELTE, Eötvös Loránd University, H - 1117 Budapest, Pázmány P. s. 1/A, Hungary
¹⁷Ewha Womans University, Seoul 120-750, Korea
¹⁸Florida Institute of Technology, Melbourne, Florida 32901, USA
¹⁹Florida State University, Tallahassee, Florida 32306, USA
²⁰Georgia State University, Atlanta, Georgia 30303, USA
²¹Hiroshima University, Kagamiyama, Higashi-Hiroshima 739-8526, Japan
²²IHEP Protvino, State Research Center of Russian Federation, Institute for High Energy Physics, Protvino, 142281, Russia
²³University of Illinois at Urbana-Champaign, Urbana, Illinois 61801, USA
²⁴Institute of Physics, Academy of Sciences of the Czech Republic, Na Slovance 2, 182 21 Prague 8, Czech Republic
²⁵Iowa State University, Ames, Iowa 50011, USA
²⁶Joint Institute for Nuclear Research, 141980 Dubna, Moscow Region, Russia
²⁷Helsinki Institute of Physics and University of Jyväskylä, P.O.Box 35, FI-40014 Jyväskylä, Finland
²⁸KEK, High Energy Accelerator Research Organization, Tsukuba, Ibaraki 305-0801, Japan
²⁹KFKI Research Institute for Particle and Nuclear Physics of the Hungarian Academy of Sciences (MTA KFKI RMKI), H-1525 Budapest 114, POBox 49, Budapest, Hungary
³⁰Korea University, Seoul 136-701, Korea
³¹Russian Research Center “Kurchatov Institute”, Moscow, Russia
³²Kyoto University, Kyoto 606-8502, Japan
³³Laboratoire Leprince-Ringuet, Ecole Polytechnique, CNRS-IN2P3, Route de Saclay, F-91128, Palaiseau, France
³⁴Lawrence Livermore National Laboratory, Livermore, California 94550, USA
³⁵Los Alamos National Laboratory, Los Alamos, New Mexico 87545, USA
³⁶LPC, Université Blaise Pascal, CNRS-IN2P3, Clermont-Fd, 63177 Aubiere Cedex, France
³⁷Department of Physics, Lund University, Box 118, SE-221 00 Lund, Sweden
³⁸University of Maryland, College Park, Maryland 20742, USA
³⁹Department of Physics, University of Massachusetts, Amherst, Massachusetts 01003-9337, USA
⁴⁰Institut für Kernphysik, University of Muenster, D-48149 Muenster, Germany
⁴¹Muhlenberg College, Allentown, Pennsylvania 18104-5586, USA
⁴²Myongji University, Yongin, Kyonggido 449-728, Korea
⁴³Nagasaki Institute of Applied Science, Nagasaki-shi, Nagasaki 851-0193, Japan
⁴⁴University of New Mexico, Albuquerque, New Mexico 87131, USA
⁴⁵New Mexico State University, Las Cruces, New Mexico 88003, USA
⁴⁶Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831, USA
⁴⁷IPN-Orsay, Université Paris Sud, CNRS-IN2P3, BPI, F-91406, Orsay, France
⁴⁸Peking University, Beijing, People's Republic of China
⁴⁹PNPI, Petersburg Nuclear Physics Institute, Gatchina, Leningrad region, 188300, Russia
⁵⁰RIKEN Nishina Center for Accelerator-Based Science, Wako, Saitama 351-0198, Japan
⁵¹RIKEN BNL Research Center, Brookhaven National Laboratory, Upton, New York 11973-5000, USA
⁵²Physics Department, Rikkyo University, 3-34-1 Nishi-Ikebukuro, Toshima, Tokyo 171-8501, Japan
⁵³Saint Petersburg State Polytechnic University, St. Petersburg, Russia
⁵⁴Universidade de São Paulo, Instituto de Física, Caixa Postal 66318, São Paulo CEP05315-970, Brazil
⁵⁵Seoul National University, Seoul 151-742, Korea

⁵⁶*Chemistry Department, Stony Brook University, Stony Brook, SUNY, New York 11794-3400, USA*⁵⁷*Department of Physics and Astronomy, Stony Brook University, SUNY, Stony Brook, New York 11794, USA*⁵⁸*University of Tennessee, Knoxville, Tennessee 37996, USA*⁵⁹*Department of Physics, Tokyo Institute of Technology, Oh-okayama, Meguro, Tokyo 152-8551, Japan*⁶⁰*Institute of Physics, University of Tsukuba, Tsukuba, Ibaraki 305, Japan*⁶¹*Vanderbilt University, Nashville, Tennessee 37235, USA*⁶²*Waseda University, Advanced Research Institute for Science and Engineering, 17 Kikui-cho, Shinjuku-ku, Tokyo 162-0044, Japan*⁶³*Weizmann Institute, Rehovot 76100, Israel*⁶⁴*Yonsei University, IPAP, Seoul 120-749, Korea*

(Received 18 June 2010; published 27 September 2010)

We have measured the azimuthal anisotropy of π^0 production for $1 < p_T < 18$ GeV/c for Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The observed anisotropy shows a gradual decrease for $3 \lesssim p_T \lesssim 7$ –10 GeV/c, but remains positive beyond 10 GeV/c. The magnitude of this anisotropy is underpredicted, up to at least ~ 10 GeV/c, by current perturbative QCD (PQCD) energy-loss model calculations. An estimate of the increase in anisotropy expected from initial-geometry modification due to gluon saturation effects and fluctuations is insufficient to account for this discrepancy. Calculations that implement a path-length dependence steeper than what is implied by current PQCD energy-loss models show reasonable agreement with the data.

DOI: 10.1103/PhysRevLett.105.142301

PACS numbers: 25.75.Dw

A central goal of high-energy nuclear physics is to understand the properties of the strongly coupled quark gluon plasma (SQGP), a new form of nuclear matter identified at the Relativistic Heavy Ion Collider (RHIC) [1]. A key tool for this goal is jet quenching or the suppression of high transverse momentum (p_T) hadron yields as a result of in-medium energy loss of high- p_T partons [2]. Such suppression was first observed in measurements of the nuclear modification factor for single hadron yield $R_{AA} = \frac{dN_{AA}}{\langle T_{AA} \rangle d\sigma_{pp}}$ [3], where dN_{AA} is the differential yield in Au + Au collisions, $d\sigma_{pp}$ is the differential cross section in $p + p$ collisions for a given p_T , and $\langle T_{AA} \rangle$ is the nuclear overlap integral for a given Au + Au centrality bin. Later on this effect was also observed in measurements of dihadron [4] and γ -hadron correlations [5].

Current theoretical descriptions of jet quenching are commonly based on a perturbative QCD (PQCD) framework [6], which assumes that the coupling of jets with the medium is weak, even though the medium itself is strongly coupled (large coupling constant α_s). Prompted by the large amount of experimental data from RHIC, several sophisticated PQCD-based models have been developed in the last decade [2,6]. These models have provided initial estimates of the properties of the SQGP, such as the momentum broadening per mean free path, $\hat{q} = \langle k_T^2 \rangle / \lambda$, and the energy loss per unit length, dE/dl [6–8].

Despite these successes, the PQCD description of jet quenching faces several challenges [9]. Besides a large discrepancy among models of extracted medium properties such as \hat{q} [8], the energy-loss models also disagree in their predictions of the azimuthal anisotropy of high p_T hadrons [8]. The latter characterizes hadron emission relative to the reaction plane (RP) angle (Ψ_{RP}), $dN/d(\phi - \Psi_{RP}) \propto (1 + 2v_2 \cos[2(\phi - \Psi_{RP})])$. Such azimuthal anisotropy ensues because the hadron yield is more suppressed along the long

axis of the almond-shaped fireball than the short axis. Thus the magnitude of the anisotropy, v_2 , is sensitive to the path-length (l) dependence of energy loss, which scales as $\Delta E \sim l$ for collisional energy loss [10], $\Delta E \sim l^2$ for coherent radiative energy loss [10], and $\Delta E \sim l^3$ for a non-perturbative energy-loss calculation using AdS/CFT gravity-gauge dual theory [11]. However, our ability to probe such l dependences hinges not only on precision data at high p_T , but also on a good understanding of the role of the initial collision geometry and space-time evolution of the medium. One geometry commonly used in energy-loss models is based on the optical Glauber model [12], which assumes a smooth Woods-Saxon nuclear geometry. Such geometry ignores the event-by-event shape distortion due to spatial fluctuations of participating nucleons [13], and a possible overall shape distortion due to gluon saturation effects, known as the CGC geometry [14]. The choice of collision geometry and medium evolution has been shown to be important for elliptic flow at low p_T [15,16], but their influences on high p_T v_2 are not well studied to date.

In this Letter we present a new measurement of the π^0 anisotropy in $\sqrt{s_{NN}} = 200$ GeV Au + Au collisions. This measurement complements our prior results [17–19], but significantly increases both the p_T reach and the statistical precision above 6 GeV/c, allowing for quantitative comparisons to energy-loss models, as well as detailed investigations of the role of the initial collision geometry.

Results were obtained from $\sim 3.5 \times 10^9$ minimum bias events taken in 2007. Event centrality was determined by the number of charged particles detected in the Beam-Beam Counters (BBC, $3.0 < |\eta| < 3.9$). A Monte Carlo (MC) Glauber model [12] was used to estimate the average number of participating nucleons (N_{part}) and $\langle T_{AA} \rangle$ for each centrality class.

Previous PHENIX analyses [19] estimated the RP using the charged particles detected in the BBC. Several new detectors, installed symmetrically on both sides of the beam line, provided additional RP measurements in 2007: the Muon Piston Calorimeters (MPC, $3.1 < |\eta| < 3.9$) and the Reaction Plane detectors in two η ranges, RXN_{in} (RXN_{out}) in $1.5(1.0) < |\eta| < 2.8(1.5)$. Each MPC is equipped with PbWO_4 crystal scintillators to detect both charged and neutral particles. Each RXN consists of 12 azimuthally segmented paddle scintillators. This analysis estimates the RP angle using both the MPC and RXN_{in} to provide good resolution, while minimizing the potential biases from jets and dijets [20]. The error on the RP angle $\Delta\Psi$, and the RP dispersion factor $\sigma_{\text{RP}} = \langle \cos 2\Delta\Psi \rangle$ are estimated by the subevent method [19], giving $\sigma_{\text{RP}} \sim 0.52$ and 0.73 in central and midcentral collisions, respectively, which is $\sim 80\%$ better than that for the BBC. The large data set and improved σ_{RP} give an equivalent of ~ 15 -fold increase in statistics over the previous measurement of v_2 [19].

The methodology for v_2 extraction follows our previous work [19]. We reconstruct the neutral pions via the $\pi^0 \rightarrow \gamma + \gamma$ decay channel with photons detected in the Electromagnetic Calorimeter (EMCal, $|\eta| < 0.35$). We apply shower shape and pair asymmetry cuts to reduce the combinatorial background. The remaining background is subtracted by the mixed event method [19]. The azimuthal distribution of the π^0 yields relative to the estimated RP angle, $\Delta\phi = \phi - \Psi_{\text{RP}}$, is divided into 6 bins in $[0, \pi/2]$ and fit to $N_0(1 + 2v_2^{\text{raw}} \cos(2\Delta\phi))$ (higher order harmonics are found to be small and do not influence v_2 value). The v_2 is then obtained by applying the dispersion correction $v_2 = v_2^{\text{raw}}/\sigma_{\text{RP}}$ for each centrality and p_T selection. The main sources of systematic uncertainties come from σ_{RP} and v_2^{raw} . The former is estimated by comparing measurements from different RP detectors, giving $\sim 10\%$

for central and peripheral collisions and $\sim 5\%$ for midcentral collisions. The latter accounts for dependence of v_2 on π^0 identification cuts, different sectors of EMCal, and different run groups, and is correlated in p_T ; it is estimated to be 10% for central collisions and 3% for other collisions.

Figures 1(a)–1(f) show $v_2(p_T)$ for six centrality bins, spanning 1–18 GeV/c. In the 10%–50% centrality range, where the signal is large and the uncertainty is small, the v_2 values above 3 GeV/c indicate a slow decrease up to 7–10 GeV/c, and remain significantly above zero at higher p_T . The ratios in Figs. 1(g)–1(i) confirm the consistency of v_2 measured using the RP from the MPC or the RXN_{in} and imply that the influence of rapidity dependent jet bias to the RP, if any, is within the statistical or systematic uncertainty of the measurement.

Figures 2(a) and 2(b) show the centrality dependence of v_2 in two high- p_T selections. They are compared with four PQCD jet-quenching model calculations, AMY, HT, and ASW from [8] and WHDG from [21].

The WHDG model was calculated for gluon density $dN_g/dy = 1000$ –1600, a range constrained by 0%–5% ($N_{\text{part}} \sim 351$) $\pi^0 R_{AA}$ data [7]; it assumes analytical Woods-Saxon nuclear geometry with a longitudinal Bjorken expansion. The AMY, HT, and ASW models were fitted independently to the 0-5% $\pi^0 R_{AA}$ data [8]; they were implemented in a 3D ideal hydrodynamic code with identical initial Wood-Saxon nuclear geometry, medium evolution and fragmentation functions. The HT and ASW models include only coherent radiative energy loss, while the AMY and WHDG also include collisional energy loss. The ASW and WHDG models predict sizable but similar v_2 , while the HT and AMY models tend to give much smaller v_2 . However, all models significantly underpredict the v_2 data in $6 < p_T < 9$ GeV/c range. For $p_T > 9$ GeV/c, ASW and WHDG results show a better agree-

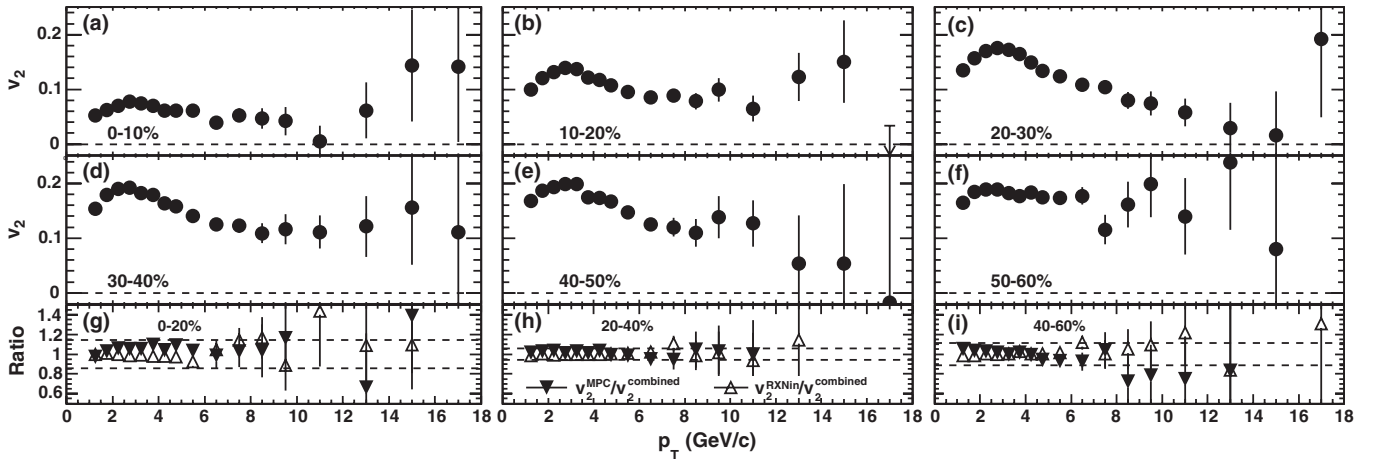


FIG. 1. (a)–(f) $\pi^0 v_2$ using combined reaction plane for MPC and RXN_{in} as a function of p_T for different centralities. (g)–(i) ratios of v_2 measured separately using MPC (solid triangles) and RXN_{in} (open triangles) to the combined result; the dashed lines indicate the systematic error. Note that the MPC and RXN_{in} are combined at the raw hit level before the RP flattening correlation [19] which unfolds for nonuniform detector acceptance; thus the combined v_2 is not a simple weighted average of v_2^{MPC} and $v_2^{\text{RXN}_{\text{in}}}$.

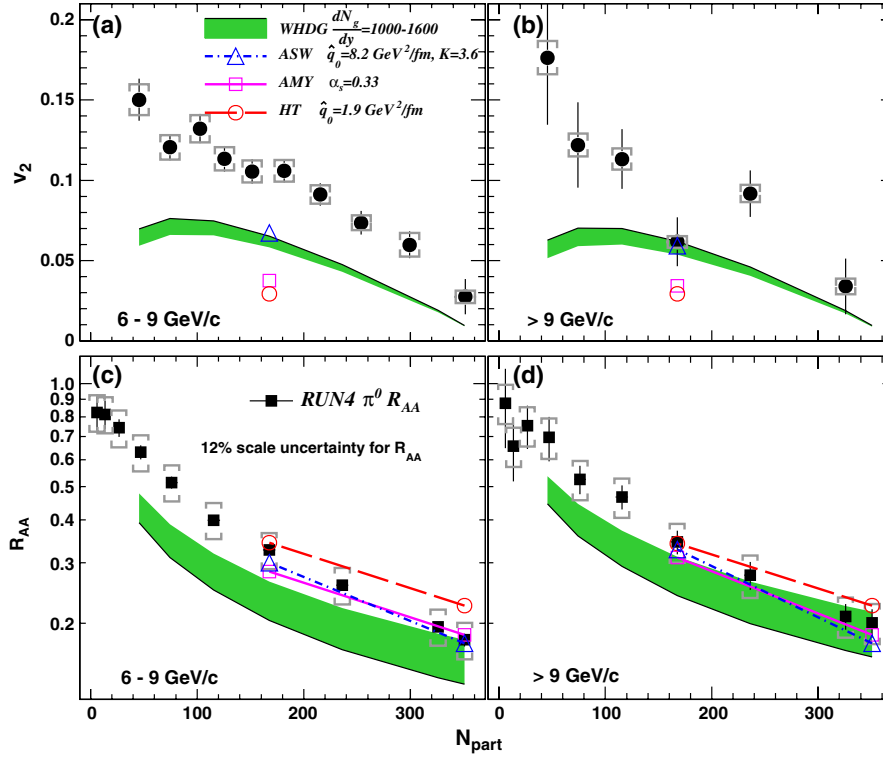


FIG. 2 (color online). (a)–(b) v_2 vs N_{part} in two p_T ranges; (c)–(d) R_{AA} vs N_{part} in same p_T ranges. Each are compared with four PQCD models from [8] (AMY, HT, ASW) with parameters quoted for quark jet at $\tau_0 = 0.6 \text{ fm}/c$, and [21] (WHDG). Log-scale is used for R_{AA} to better visualize model calculations. Note that the $\frac{dN_g}{dy} = 1000$ of WHDG corresponds to lower (upper) boundary of the shaded bands for v_2 (R_{AA}), while $\frac{dN_g}{dy} = 1600$ corresponds to upper (lower) boundary for v_2 (R_{AA}).

ment with the 20%–30% ($N_{part} \sim 167$) centrality bin due to a slow decrease of v_2 with p_T [see Fig. 1(b)]. This is accidental, because the v_2 values for the other centrality bins remain large, and are significantly above the WHDG calculations (the p value for the agreement is $< 10^{-4}$).

In all these models, the inclusive suppression R_{AA} and v_2 are anticorrelated; i.e., a smaller R_{AA} implies a larger v_2 and vice versa. Consequently, more information can be obtained by comparing the data with a given model for both R_{AA} and v_2 . Figures 2(c) and 2(d) compare the centrality dependence of $\pi^0 R_{AA}$ data to four model calculations for the same two p_T ranges [22]. The calculations are available for a broad centrality range for WHDG, but only in 0%–5% and 20%–30% centrality bins for AMY, HT and ASW. The level of agreement varies among the models. The HT calculations are slightly above the data in the most central bin, while WHDG systematically underpredicts the data over the full centrality range, though better agreement with the data is obtained for $p_T > 9 \text{ GeV}/c$. On the other hand, ASW and AMY calculations agree with the data very well in both p_T ranges. The different levels of agreement among the models are partially due to their different trends of R_{AA} with p_T : WHDG and ASW results have stronger p_T dependences than what is found in the data, and tend to deviate at low p_T when fitted to the full p_T range [7,8].

Given the larger fractional systematic error for R_{AA} measurements compared to the v_2 measurements, the deviation of $v_2(N_{part})$ from the data is more dramatic than that for the $R_{AA}(N_{part})$. Nevertheless, Fig. 2 clearly shows the importance for any model to simultaneously describe the R_{AA} and the azimuthal anisotropy of the data.

The fact that the high p_T v_2 at RHIC exceeds expectations of PQCD jet-quenching models was first pointed out in Ref. [23] in 2002. This was not a serious issue back then since the p_T reach of early measurements was rather limited, and the v_2 could be strongly influenced, up to 6 GeV/c for pions, by collective flow and recombination effects rather than jet quenching [24]. Figure 2 clearly shows that the v_2 at p_T above 6 or even 9 GeV/c still exceeds the PQCD-based energy-loss models. It is possible that geometrical effects due to fluctuations and CGC effects, ignored in these models, can increase the calculated v_2 ; it is also possible that the energy-loss process in the SQGP has a steeper l dependence (e.g., AdS/CFT) than what is currently implemented in these models.

To test whether these two ideas could bridge the difference between data and theory, we compare the data with the JW model from [25]. This model is based on a naïve jet absorption picture with an exponential survival probability $e^{-\kappa I}$ for jets, where the line integral $I = \int dl \rho$ is chosen

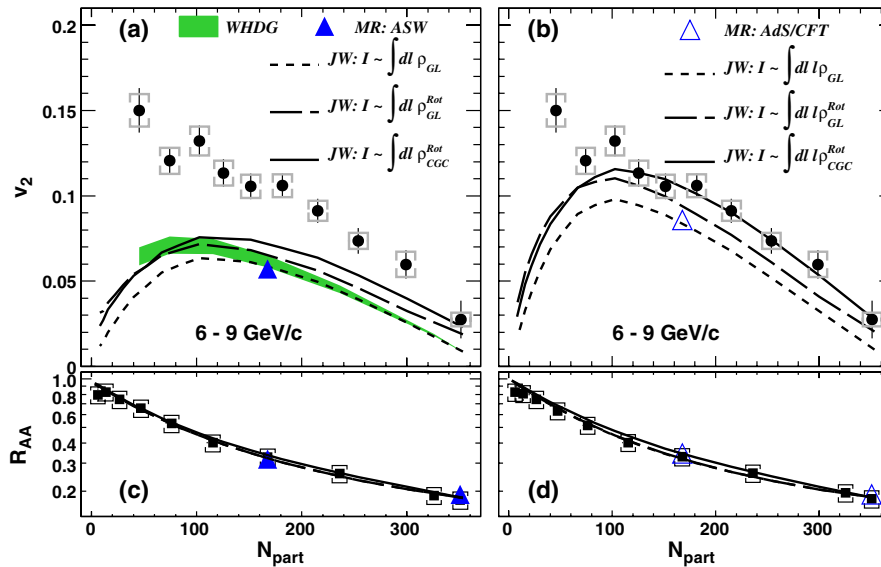


FIG. 3 (color online). v_2 vs N_{part} in 6–9 GeV/c compared with various models: (a) WHDG [21] (shaded bands), ASW [27] (solid triangle), and three JW calculations [25] with quadratic l dependence with longitudinal expansion for Glauber geometry (dashed lines), rotated Glauber geometry (long dashed lines), and rotated CGC geometry (solid lines); (b) Same as (a) except that AdS/CFT modified calculation in ASW framework (triangle) from [27] is shown and the JW calculations were done for cubic l dependence with longitudinal expansion; (c)–(d) the comparison of calculated R_{AA} s from these models with data.

for a quadratic dependence ($\sim ldl$) of energy loss in a longitudinally expanding medium ($\sim 1/l$), and κ is tuned to reproduce the central R_{AA} data. The medium density ρ is given by two leading candidates of the initial geometry: MC Glauber geometry $\rho_{\text{GL}}(x, y) = 0.43\rho_{\text{part}}(x, y) + 0.14\rho_{\text{coll}}(x, y)$, i.e., a mixture of participant density profile and binary collision profile from PHOBOS [26]; and MC CGC geometry $\rho_{\text{CGC}}(x, y)$ of Drescher and Nara [14]. The effect of fluctuations for both profiles were included via the standard rotation procedure [13]. The short-dashed curves in Fig. 3(a) show that the result for Glauber geometry without rotation (ρ_{GL}) compares reasonably well with those from WHDG [21] and a version of ASW model from [27]. Consequently, we use the JW model to estimate the shape distortions due to fluctuations and CGC effects. The results for Glauber geometry with rotation ($\rho_{\text{GL}}^{\text{Rot}}$) and CGC geometry with rotation ($\rho_{\text{CGC}}^{\text{Rot}}$) each lead to a $\sim 15\%$ – 20% increase of v_2 in midcentral collisions. However, these calculated results still fall below the data.

Figure 3(b) compares the same data with three JW models for the same matter profiles, but calculated for a line integral motivated by AdS/CFT correspondence $I = \int dl l \rho$. The stronger l dependence for ρ_{GL} significantly increases (by $>50\%$) the calculated v_2 , and brings it close to the data for midcentral collisions. However, a sizable fractional difference in the central bin seems to require an additional increase from fluctuations and CGC geometry. Figure 3(b) also shows a MR model from [27], which implements the AdS/CFT l dependence within the ASW framework [28]; it compares reasonably well with the JW

model for ρ_{GL} (short-dashed curves). Note that the MR and JW models in Fig. 3 have been tuned independently to reproduce the 0–5% π^0 R_{AA} data, and they all describe the centrality dependence of R_{AA} very well [see Figs. 3(c) and 3(d)]. On the other hand, these models predict a stronger suppression for dihadrons than for single hadrons, opposite to experimental findings [29]; thus a global confrontation of any model with all experimental observables is warranted.

In summary, we presented results on π^0 azimuthal anisotropy (v_2) in $1 < p_T < 18$ GeV/c in Au + Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The measurements indicate sizable $v_2(p_T)$ that decreases gradually for $3 \lesssim p_T \lesssim 7$ – 10 GeV/c, but remains positive for $p_T > 10$ GeV/c. This large v_2 exceeds expectations of PQCD energy-loss models even at $p_T \sim 10$ GeV/c. Estimates of the v_2 increase due to modifications of initial geometry from gluon saturation effects and fluctuations indicate that they are insufficient to reconcile data and theory. Incorporating an AdS/CFT-like path-length dependence for jet quenching in a PQCD-based framework [27] and a schematic model [25] both compare well with the data. However, more detailed study beyond these simplified models are required to quantify the nature of the path-length dependence. Our precision data provide key constraints on the initial geometry, medium space-time evolution, and the jet-quenching mechanisms.

We thank the staff of the Collider-Accelerator and Physics Departments at BNL for their vital contributions. We acknowledge support from the Office of Nuclear

Physics in DOE Office of Science and NSF (USA), MEXT and JSPS (Japan), CNPq and FAPESP (Brazil), NSFC (China), MSMT (Czech Republic), IN2P3/CNRS and CEA (France), BMBF, DAAD, and AvH (Germany), OTKA (Hungary), DAE and DST (India), ISF (Israel), NRF and WCU (Korea), MES, RAS, and FAAE (Russia), VR and KAW (Sweden), U.S. CRDF for the FSU, US-Hungary Fulbright, and US-Israel BSF.

*Deceased.

†PHENIX Spokesperson.

jacak@skipper.physics.sunysb.edu

- [1] K. Adcox *et al.*, *Nucl. Phys.* **A757**, 184 (2005); J. Adams *et al.*, *Nucl. Phys.* **A757**, 102 (2005); B. B. Back *et al.*, *Nucl. Phys.* **A757**, 28 (2005); I. Arsene *et al.*, *Nucl. Phys.* **A757**, 1 (2005).
- [2] M. Gyulassy, I. Vitev, X.N. Wang, and B.W. Zhang, [arXiv:nucl-th/0302077](https://arxiv.org/abs/nucl-th/0302077).
- [3] K. Adcox *et al.*, *Phys. Rev. Lett.* **88**, 022301 (2001).
- [4] C. Adler *et al.*, *Phys. Rev. Lett.* **90**, 082302 (2003).
- [5] A. Adare *et al.*, *Phys. Rev. C* **80**, 024908 (2009).
- [6] A. Majumder and M. Van Leeuwen, [arXiv:1002.2206](https://arxiv.org/abs/1002.2206).
- [7] A. Adare *et al.*, *Phys. Rev. C* **77**, 064907 (2008).
- [8] S. A. Bass *et al.*, *Phys. Rev. C* **79**, 024901 (2009).
- [9] B. Muller, *Prog. Theor. Phys. Suppl.* **174**, 103 (2008).
- [10] S. Peigne and A. V. Smilga, *Phys. Usp.* **52**, 659 (2009).
- [11] S. S. Gubser, D. R. Gulotta, S. S. Pufu, and F. D. Rocha, *J. High Energy Phys.* **10** (2008) 052; F. Dominguez, C. Marquet, A. H. Mueller, B. Wu, and B. W. Xiao, *Nucl. Phys.* **A811**, 197 (2008).
- [12] M. L. Miller, K. Reygers, S. J. Sanders, and P. Steinberg, *Annu. Rev. Nucl. Part. Sci.* **57**, 205 (2007).
- [13] B. Alver *et al.*, *Phys. Rev. C* **77**, 014906 (2008).
- [14] H. J. Drescher and Y. Nara, *Phys. Rev. C* **75**, 034905 (2007).
- [15] M. Luzum and P. Romatschke, *Phys. Rev. C* **78**, 034915 (2008); **79**, 039903(E) (2009).
- [16] T. Hirano and Y. Nara, *Phys. Rev. C* **79**, 064904 (2009).
- [17] S. S. Adler *et al.*, *Phys. Rev. Lett.* **96**, 032302 (2006).
- [18] S. S. Adler *et al.*, *Phys. Rev. C* **76**, 034904 (2007).
- [19] S. Afanasiev *et al.*, *Phys. Rev. C* **80**, 054907 (2009).
- [20] A. Adare *et al.*, *Phys. Rev. C* **78**, 014901 (2008).
- [21] S. Wicks, W. Horowitz, M. Djordjevic, and M. Gyulassy, *Nucl. Phys.* **A784**, 426 (2007);
- [22] A. Adare *et al.*, *Phys. Rev. Lett.* **101**, 232301 (2008).
- [23] E. V. Shuryak, *Phys. Rev. C* **66**, 027902 (2002).
- [24] V. Greco, C. M. Ko, and P. Levai, *Phys. Rev. C* **68**, 034904 (2003); R. J. Fries, B. Muller, C. Nonaka, and S. A. Bass, *Phys. Rev. C* **68**, 044902 (2003).
- [25] A. Drees, H. Feng, and J. Jia, *Phys. Rev. C* **71**, 034909 (2005); J. Jia and R. Wei, *Phys. Rev. C* **82**, 024902 (2010).
- [26] B. B. Back *et al.*, *Phys. Rev. C* **65**, 061901 (2002).
- [27] C. Marquet and T. Renk, *Phys. Lett. B* **685**, 270 (2010).
- [28] The two MR models (ASW and AdS/CFT) in Fig. 3 are based on a 3D ideal hydrodynamic code slightly different from that of the ASW model shown in Fig. 2.
- [29] A. Adare *et al.*, *Phys. Rev. Lett.* **104**, 252301 (2010).